

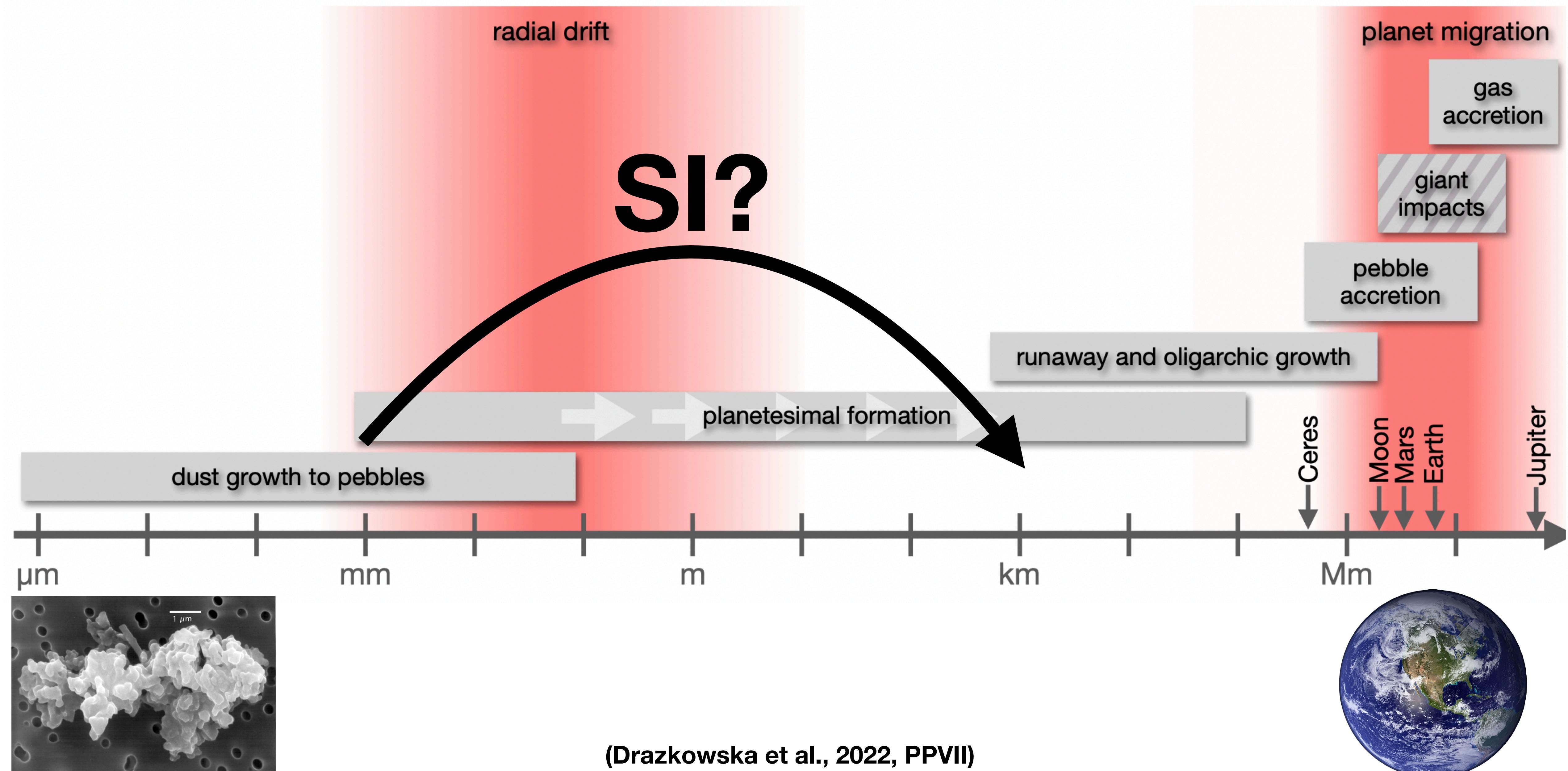
Streaming Instability: Reborn

Min-Kai Lin

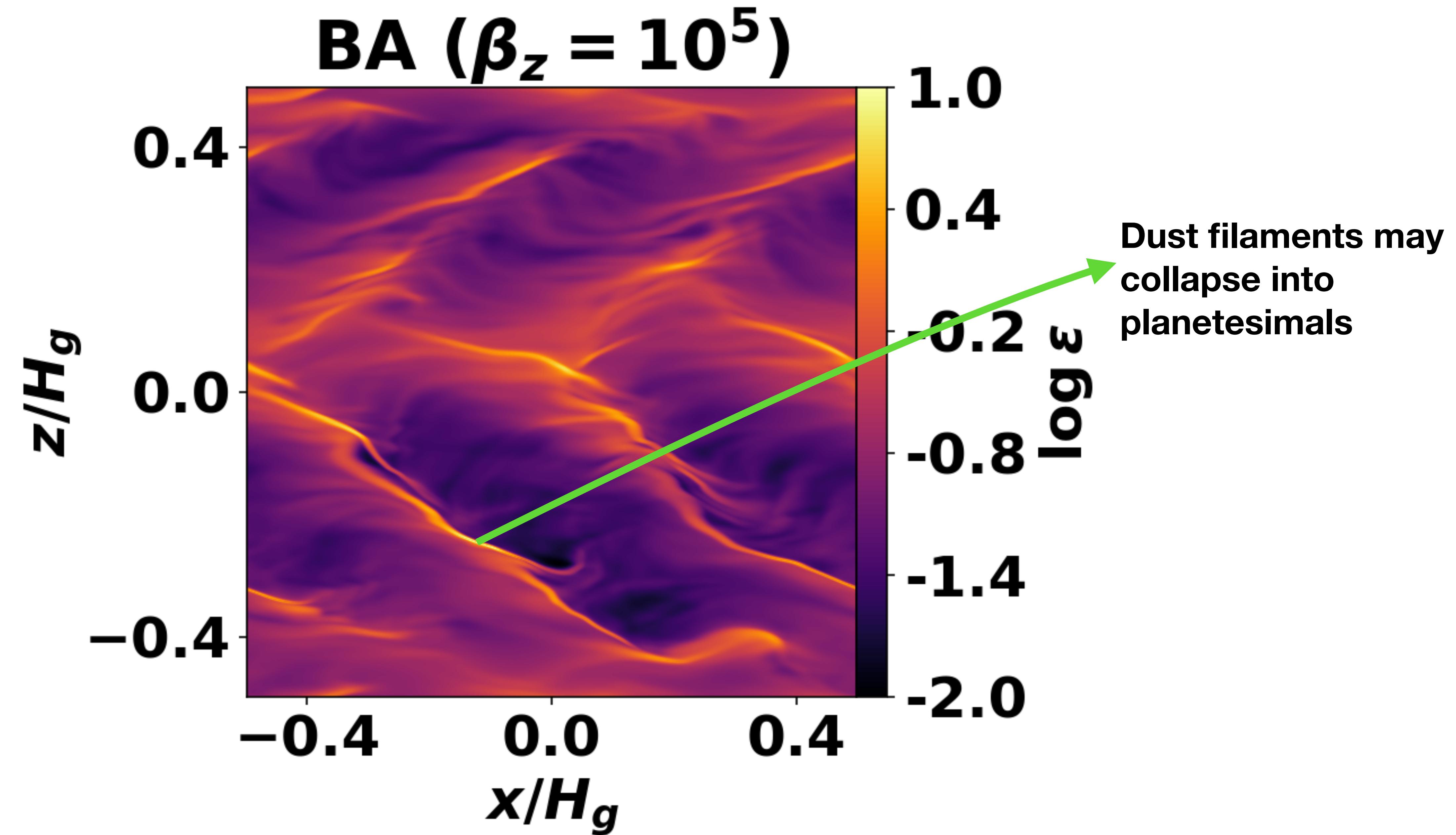
December 2023



Streaming instability for planetesimal formation



Streaming instability of dusty gas



The SI is both simple and complex

Abstract interpretation

- Resonance between dust-gas drift and inertial waves (Squire & Hopkins 2018)

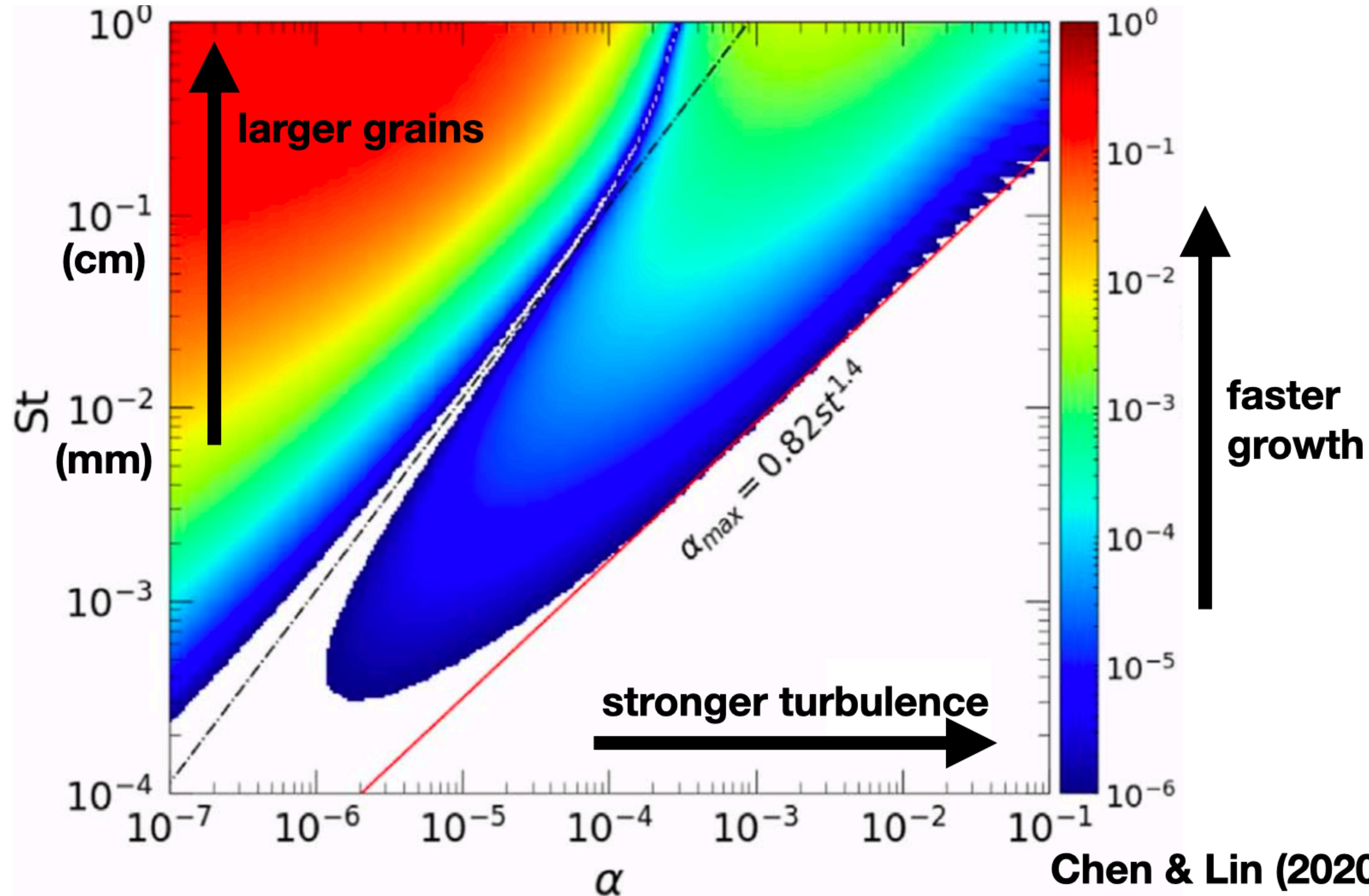
Simple ingredients

- Mutually interacting dust and gas in rotation + relative drift
- But PPDs are much more

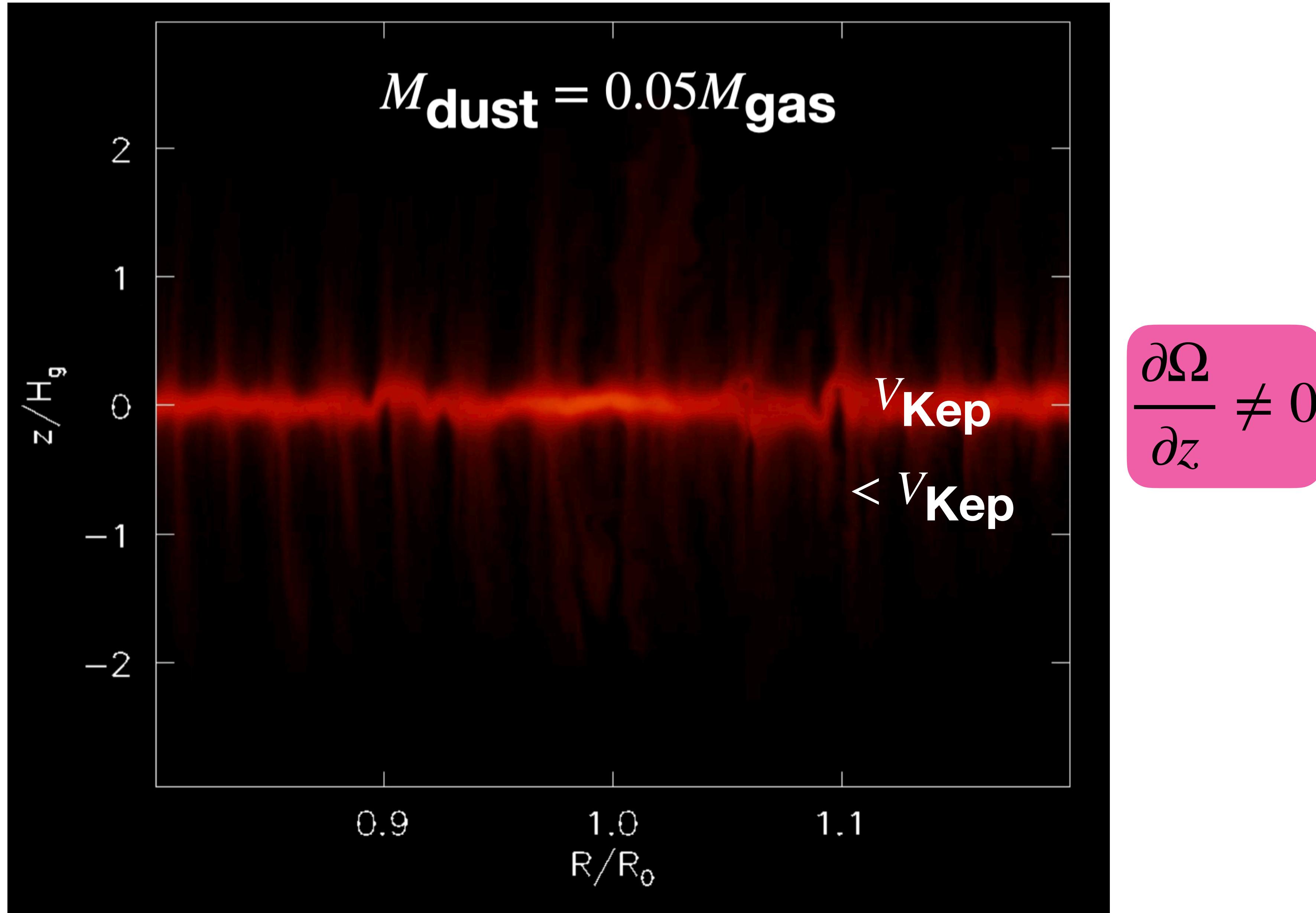
Extensions to the SI

- turbulence → **Chen & Lin (2020)**
- vertical structure → **Lin (2021)**
- magnetic fields → **Lin & Hsu (2022), Hsu & Lin (2022),
Wu, Lin et al. (accepted)**
- thermodynamics → **Lehmann & Lin (2023)**

Streaming instability is easily killed by turbulent viscosity

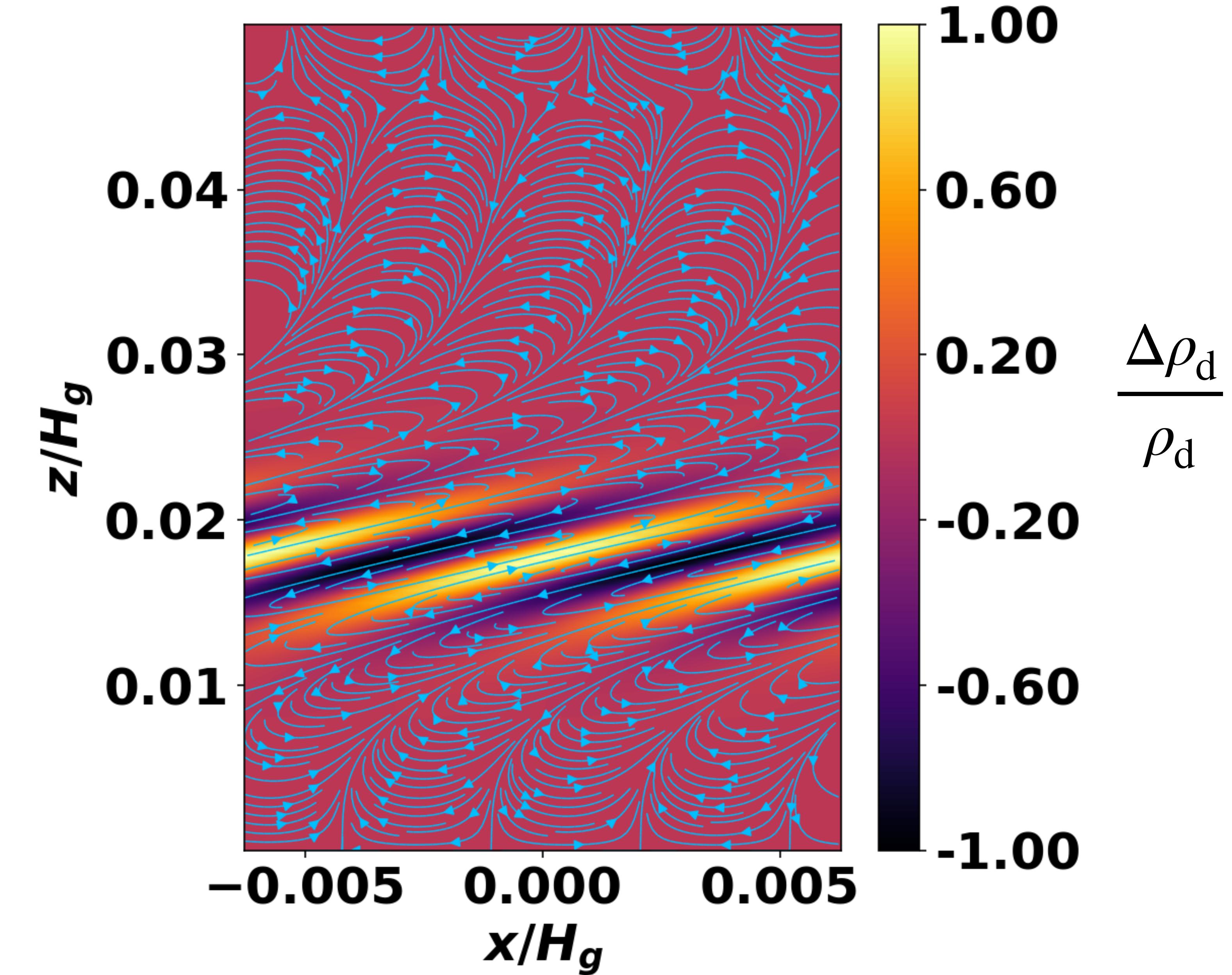


Stratified dust layers

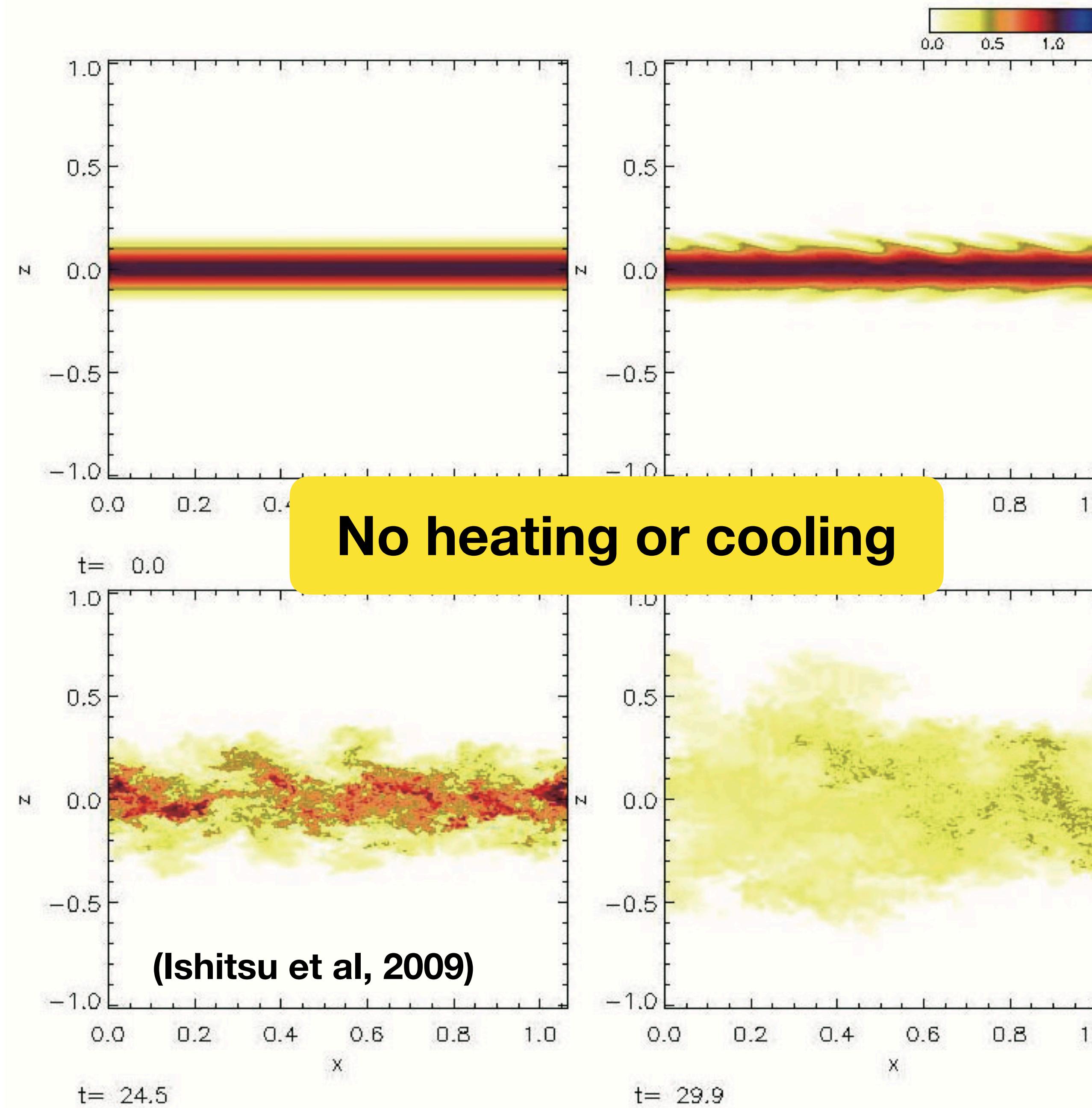


“Vertically shearing SI” in stratified disks

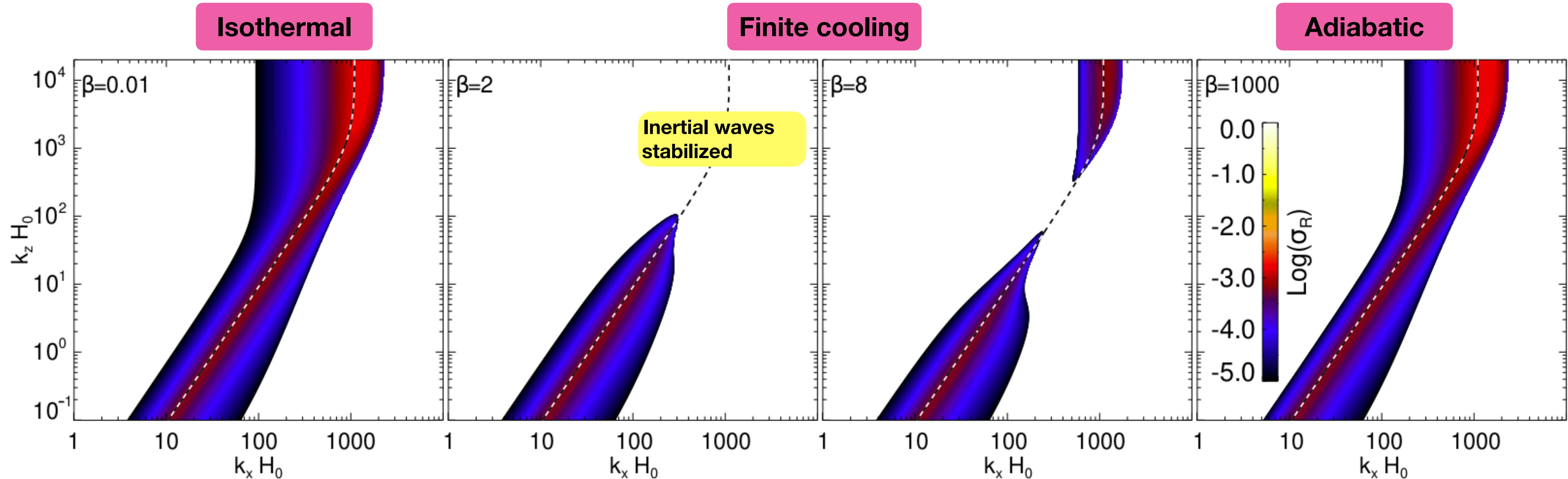
$$S_{\text{grow}} \sim \Omega$$



Vertically shearing SIs grow fast but...



SI + radial buoyancy $N_R^2 > 0$



So far, not so good

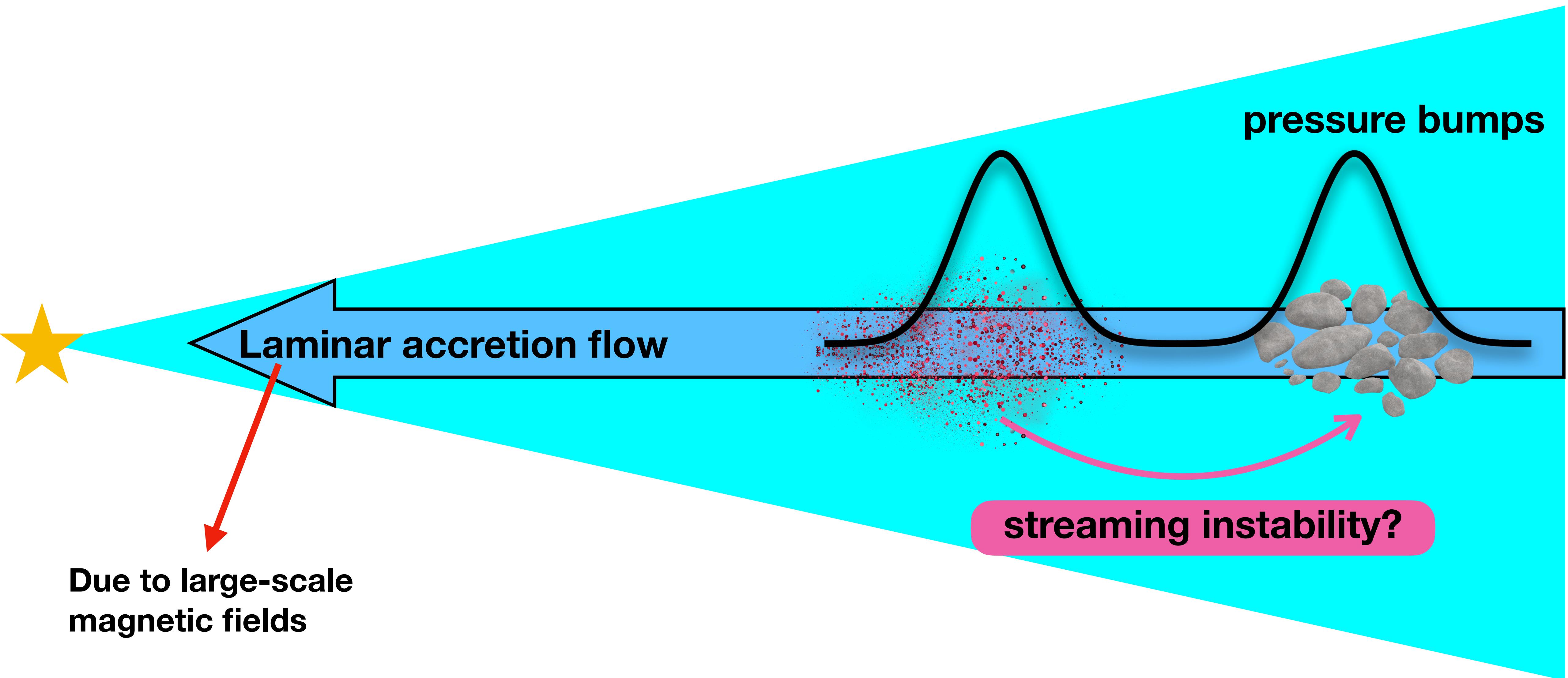
- turbulence
- vertical structure
- thermodynamics



SI weakened

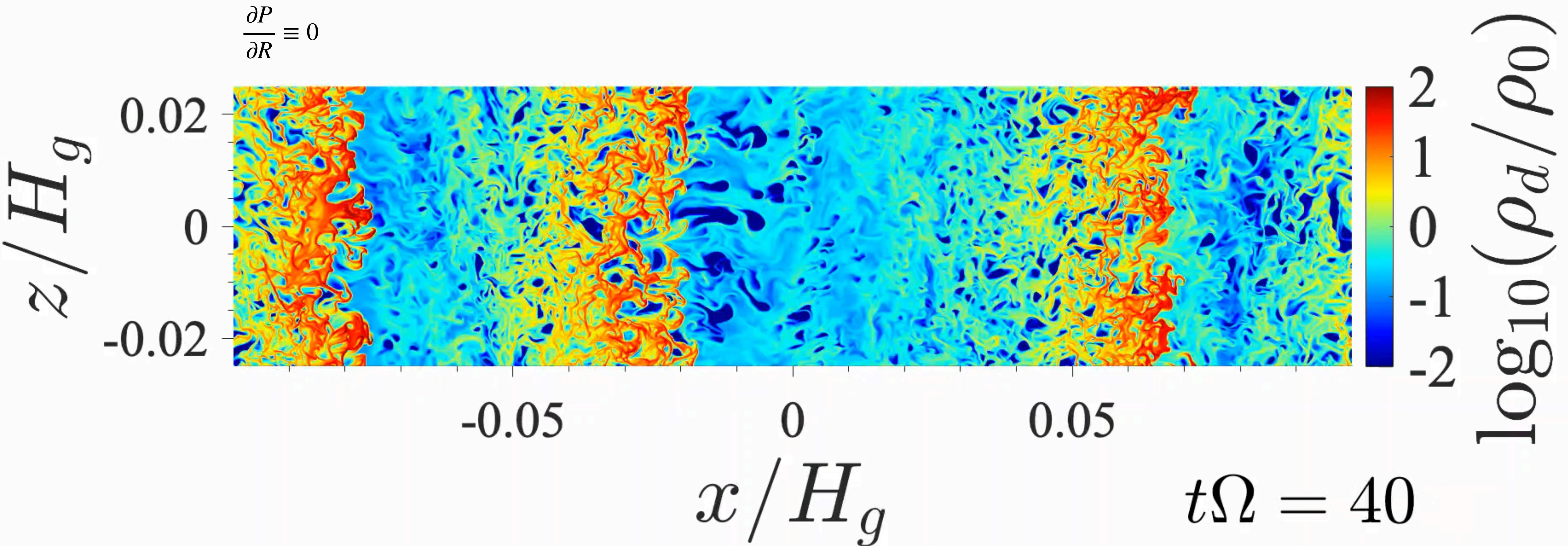


What about magnetic fields?

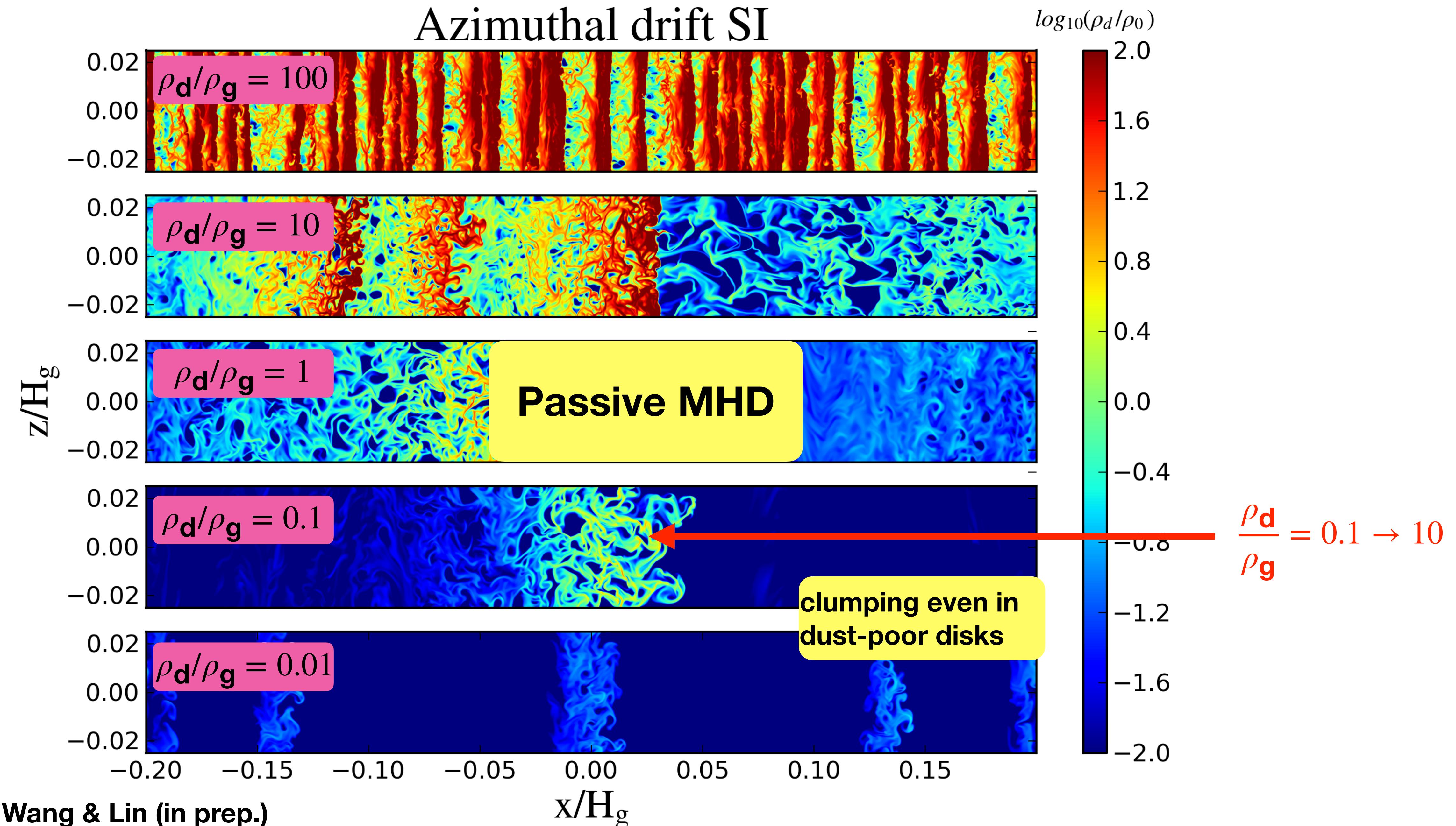


(e.g. Riols et al. 2020, Cui & Bai 2021)

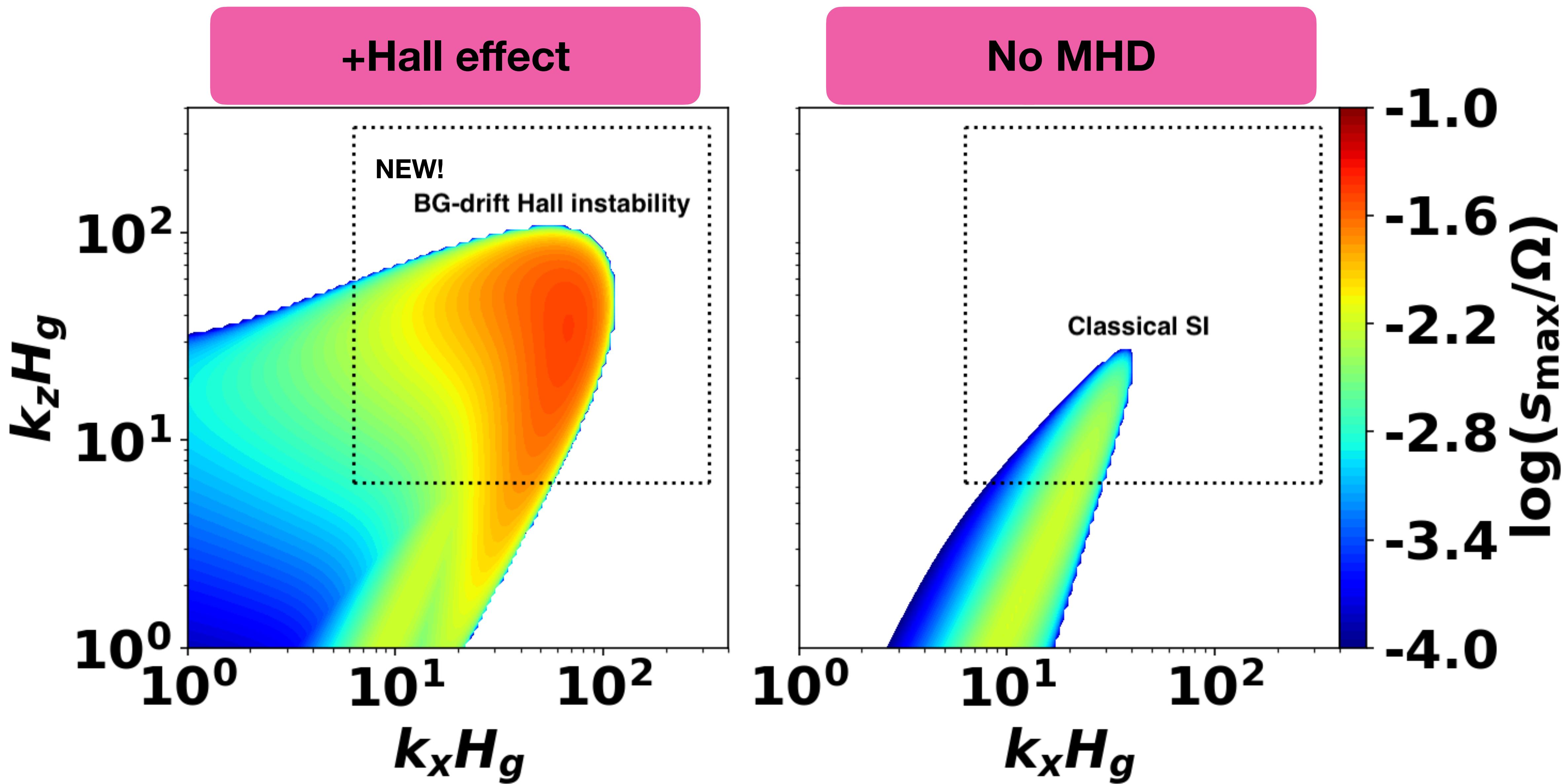
Nonlinear evolution of the SI in accreting disks



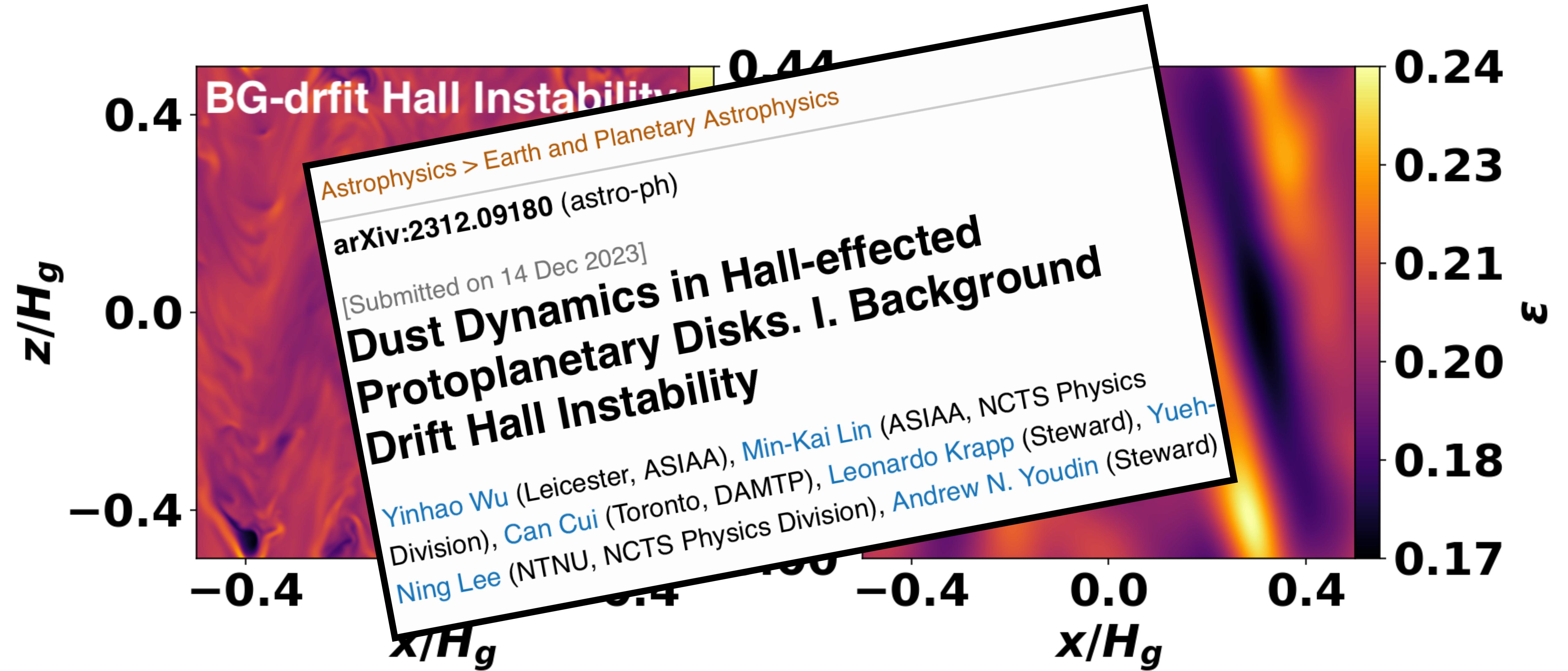
Parameter study



Live MHD & dust dynamics



Spectral simulations with Dedalus



Summary

- **Kill the SI:** turbulence, vertical structure, thermodynamics
- **Revive the SI:** gas accretion, magnetic fields
- **Future:** convection, coagulation, improved simulations, etc.

Thank you
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