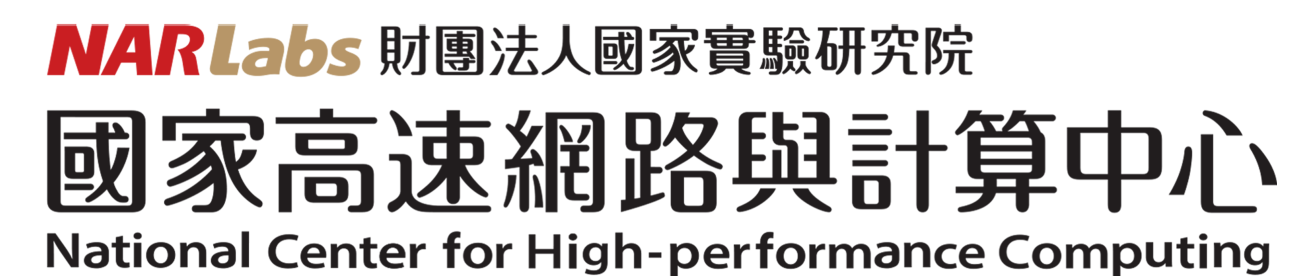


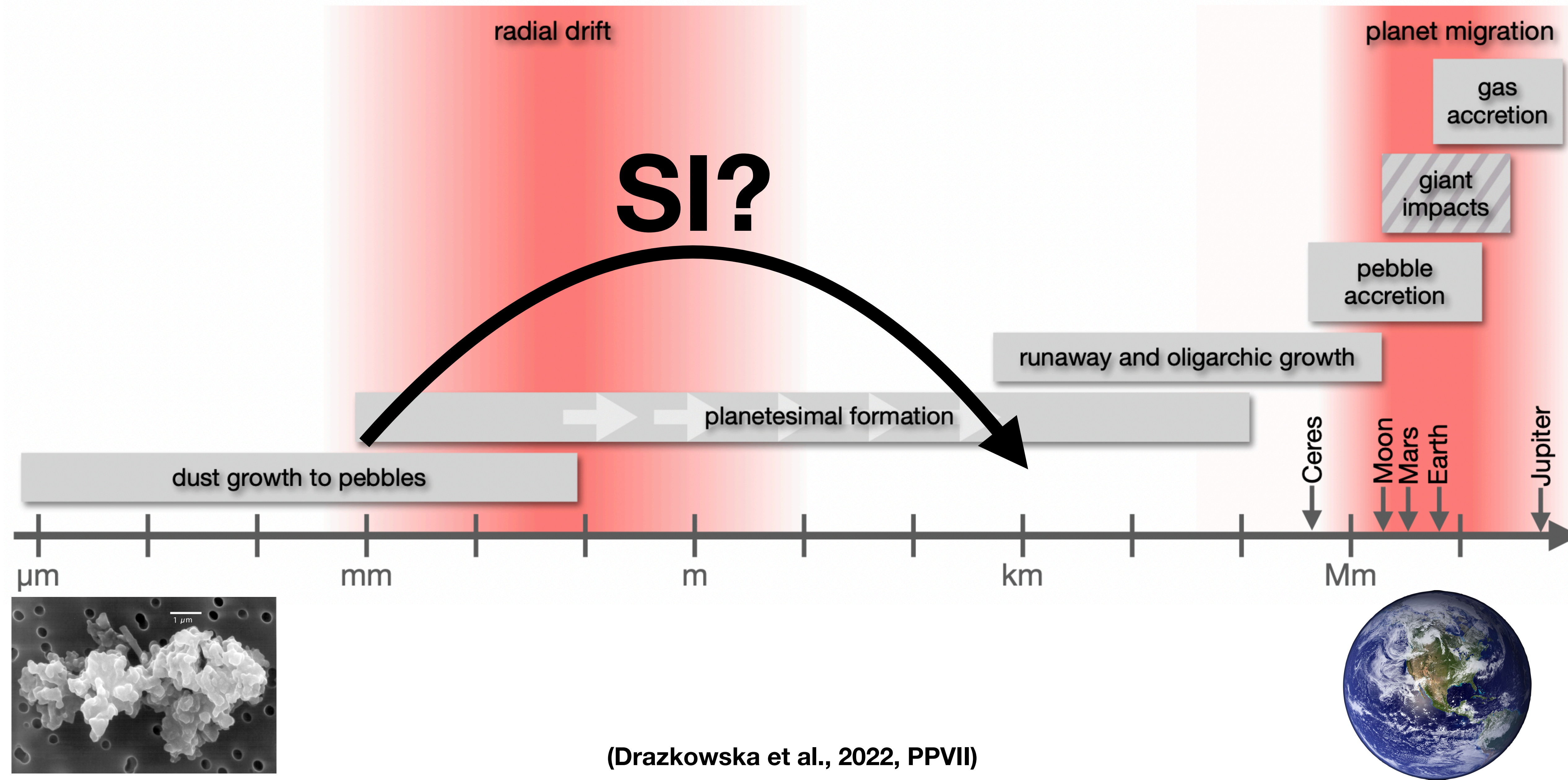
Streaming Instability: Reborn

Min-Kai Lin

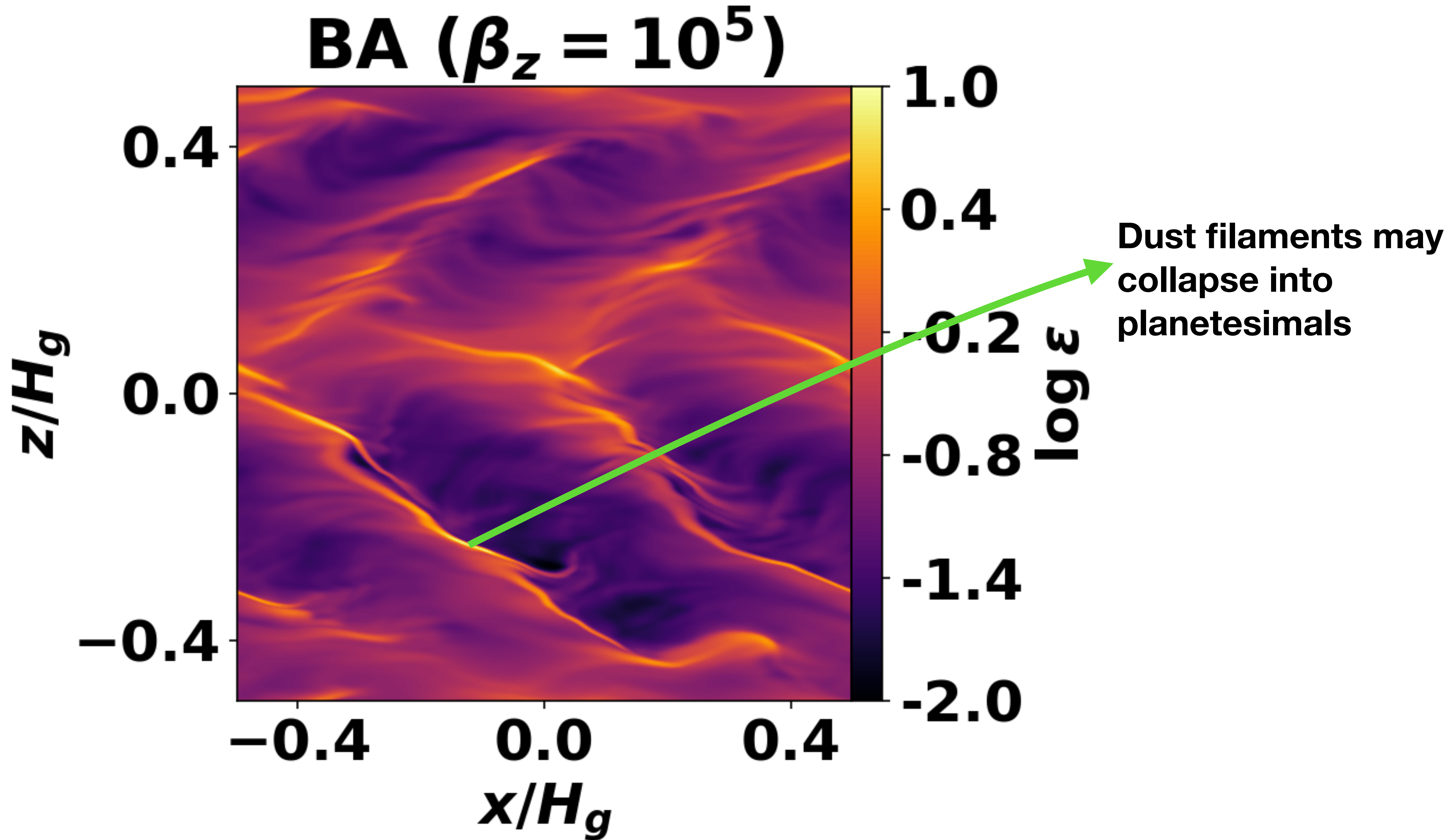
December 2023



Streaming instability for planetesimal formation



Streaming instability of dusty gas



The SI is both simple and complex

Abstract interpretation

- **Resonance between dust-gas drift and inertial waves (Squire & Hopkins 2018)**

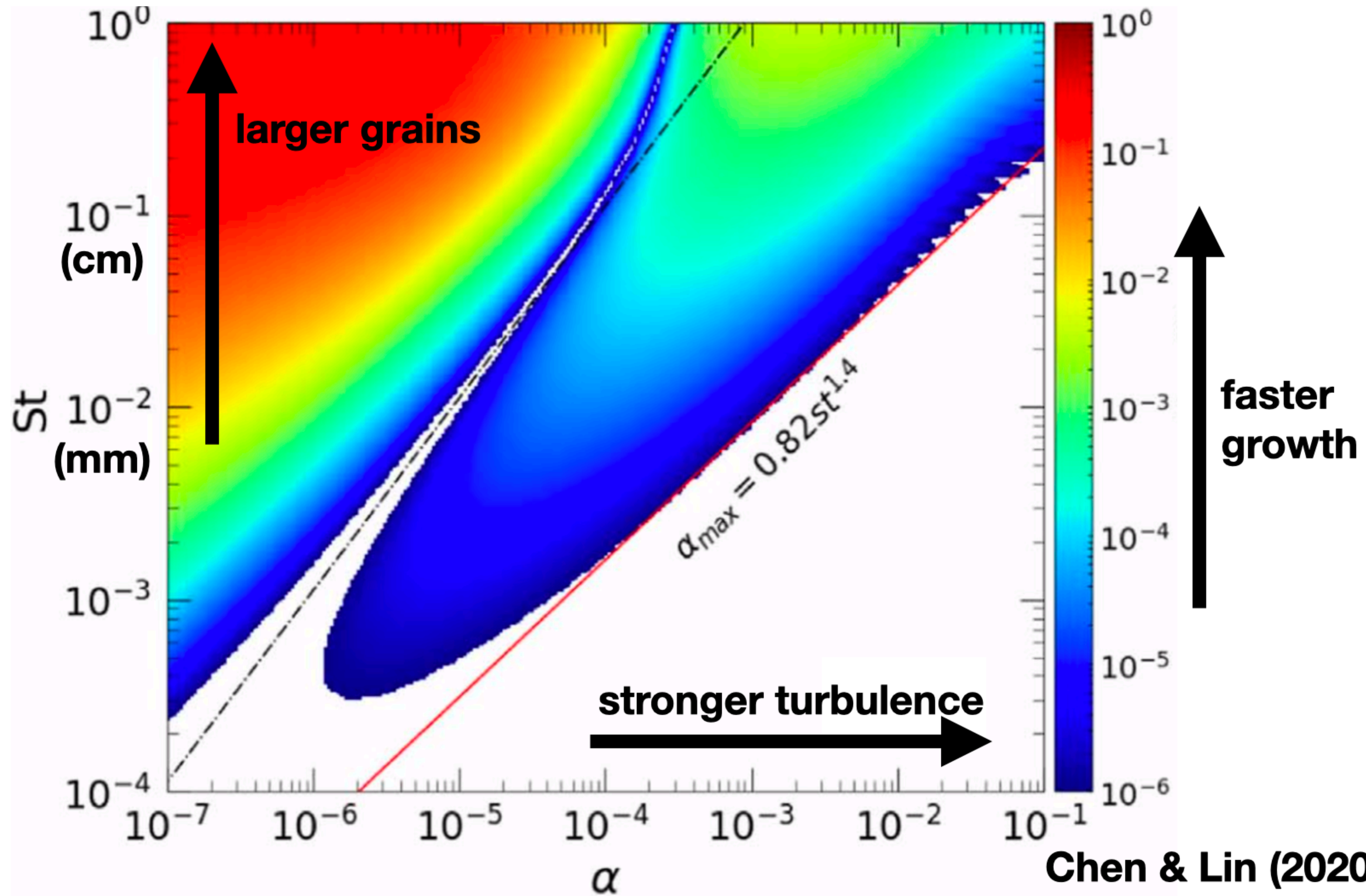
Simple ingredients

- **Mutually interacting dust and gas in rotation + relative drift**
- **But PPDs are much more**

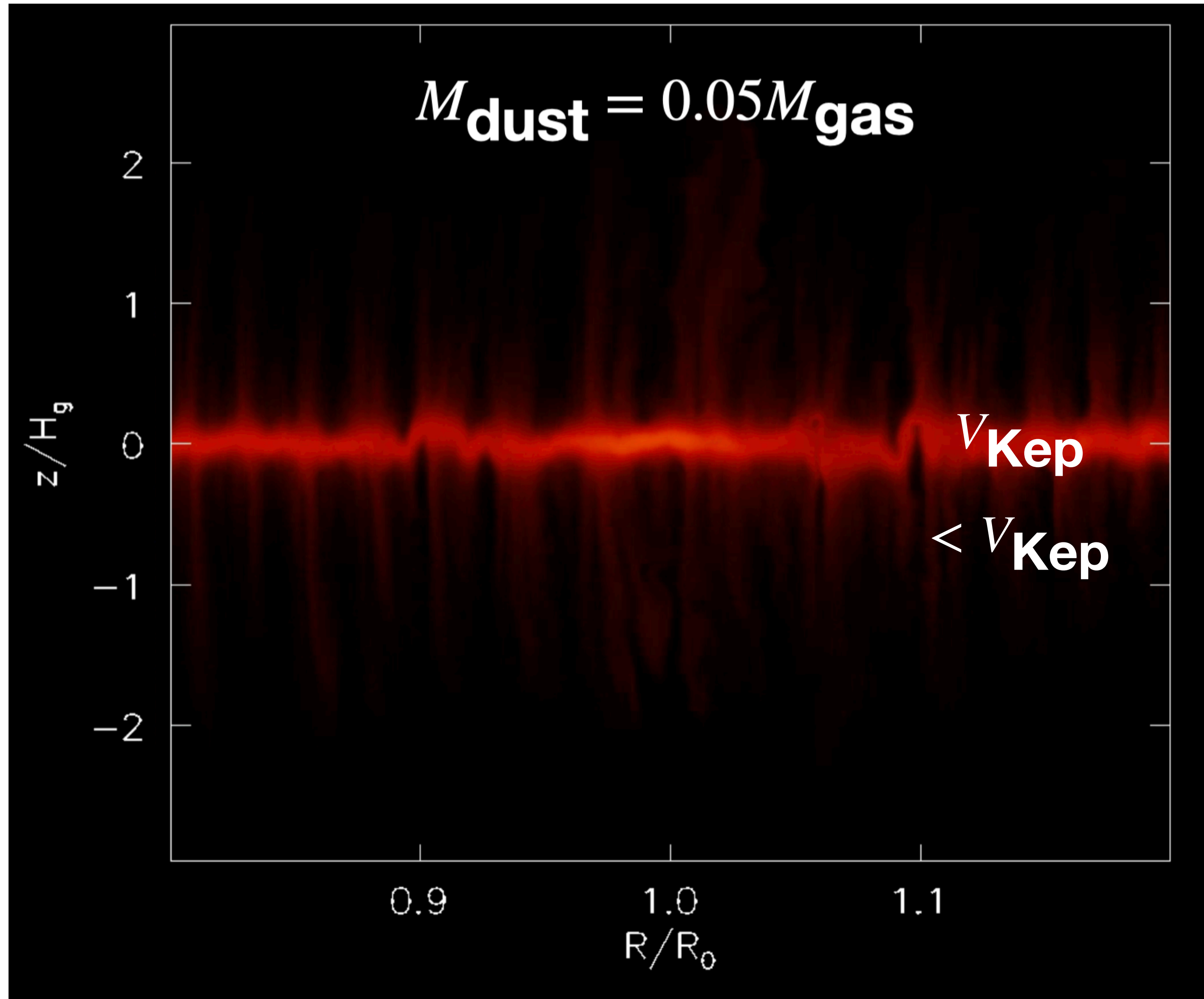
Extensions to the SI

- **turbulence** → **Chen & Lin (2020)**
- **vertical structure** → **Lin (2021)**
- **magnetic fields** → **Lin & Hsu (2022), Hsu & Lin (2022),
Wu, Lin et al. (accepted)**
- **thermodynamics** → **Lehmann & Lin (2023)**

Streaming instability is easily killed by turbulent viscosity



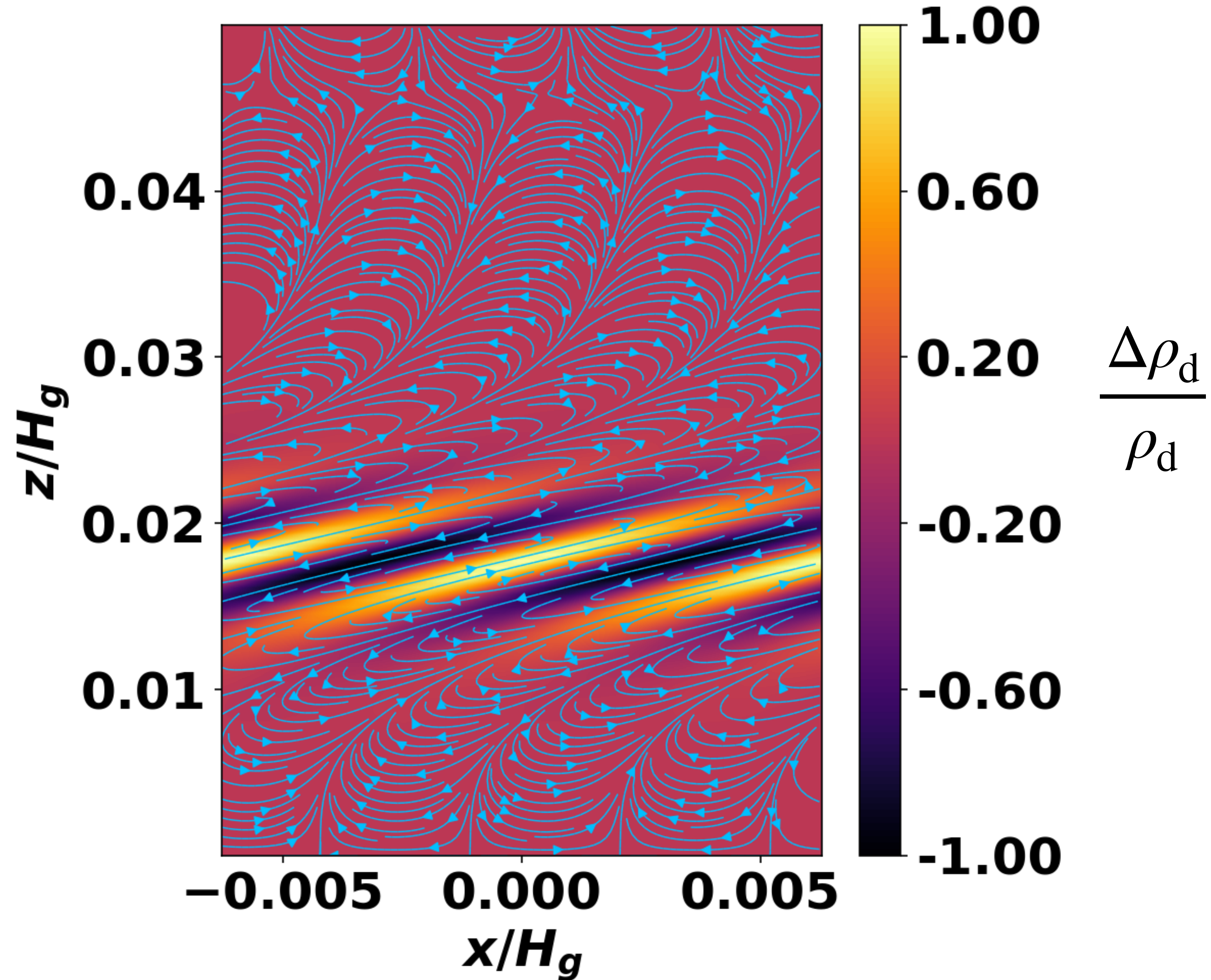
Stratified dust layers



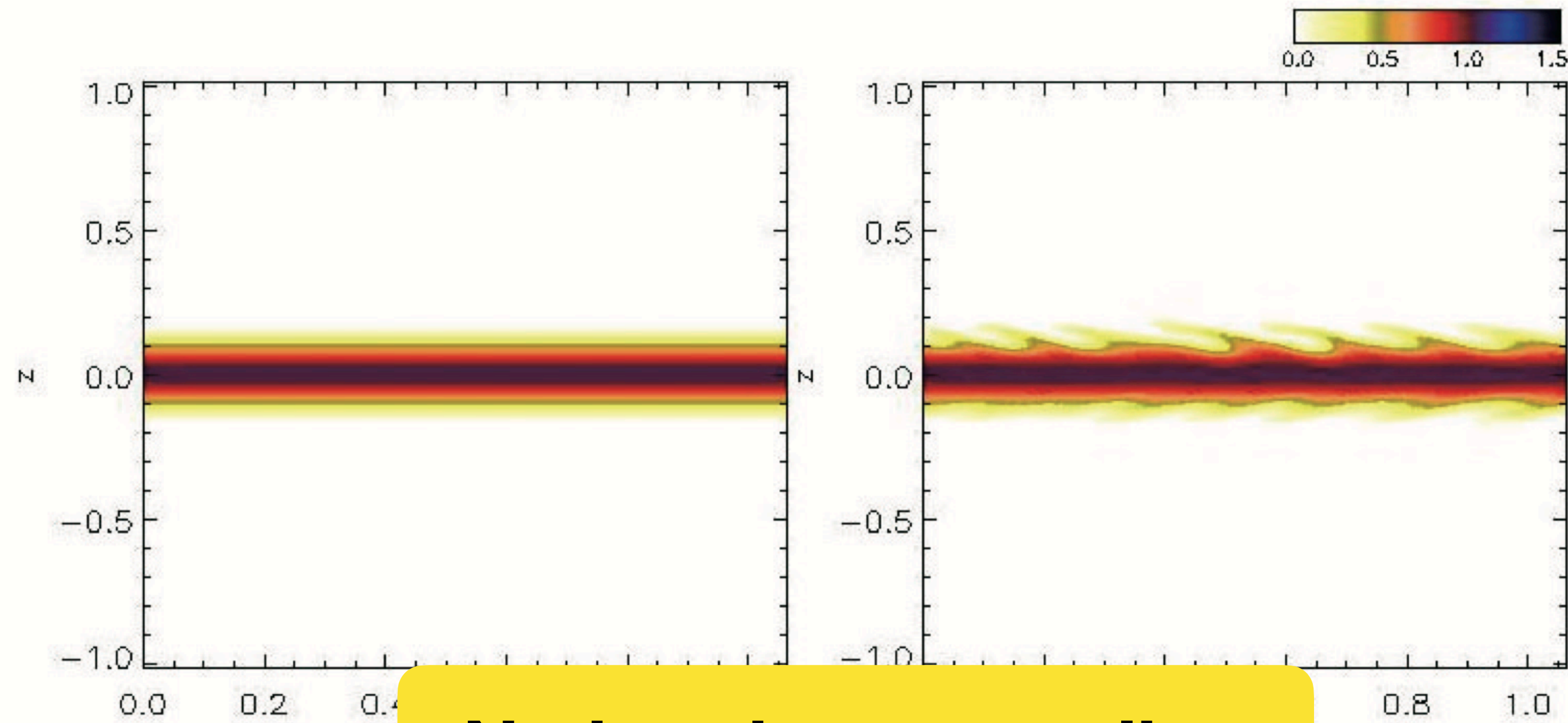
$$\frac{\partial \Omega}{\partial z} \neq 0$$

“Vertically shearing SI” in stratified disks

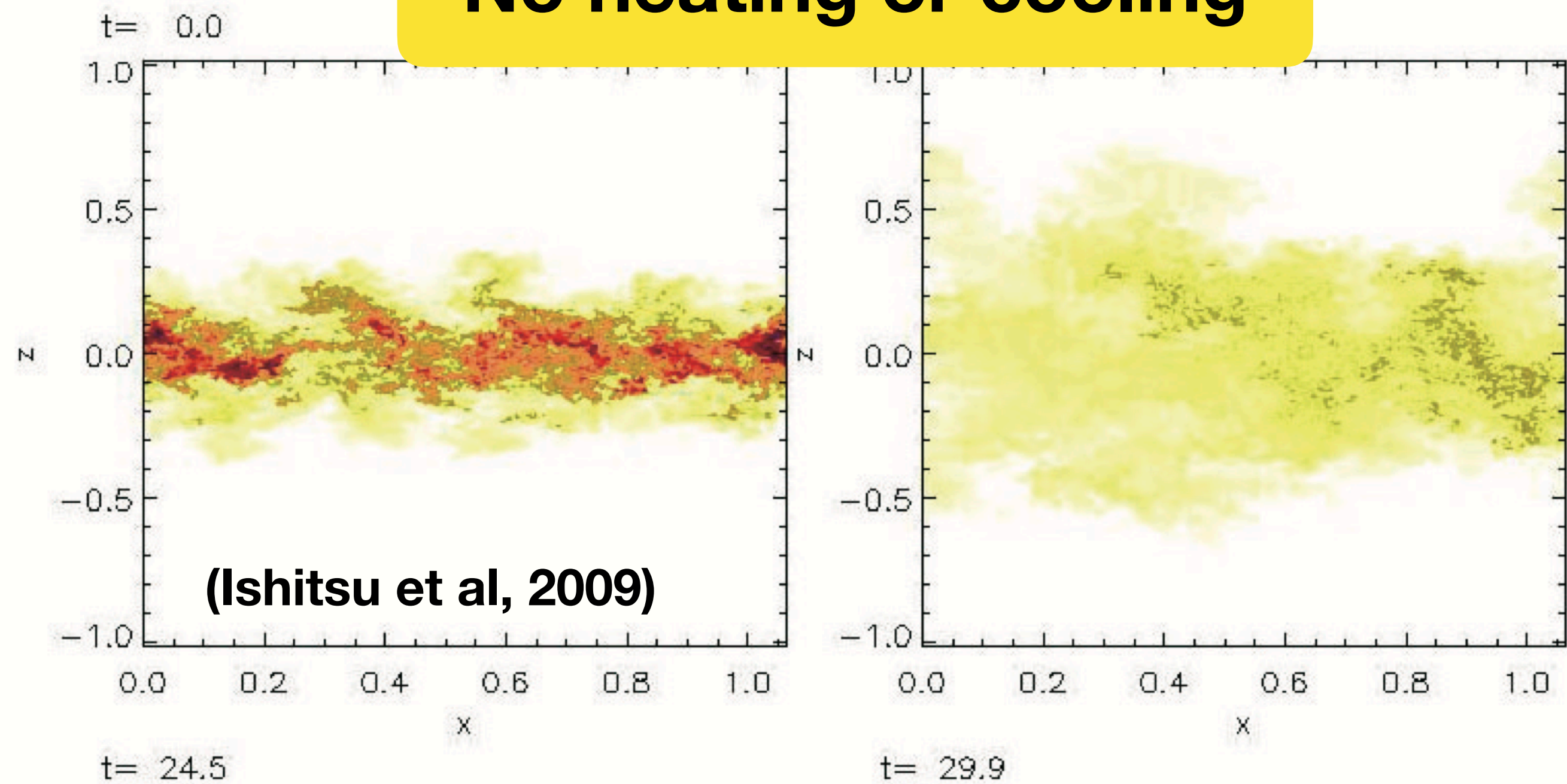
$$S_{\text{grow}} \sim \Omega$$



Vertically shearing SIs grow fast but...



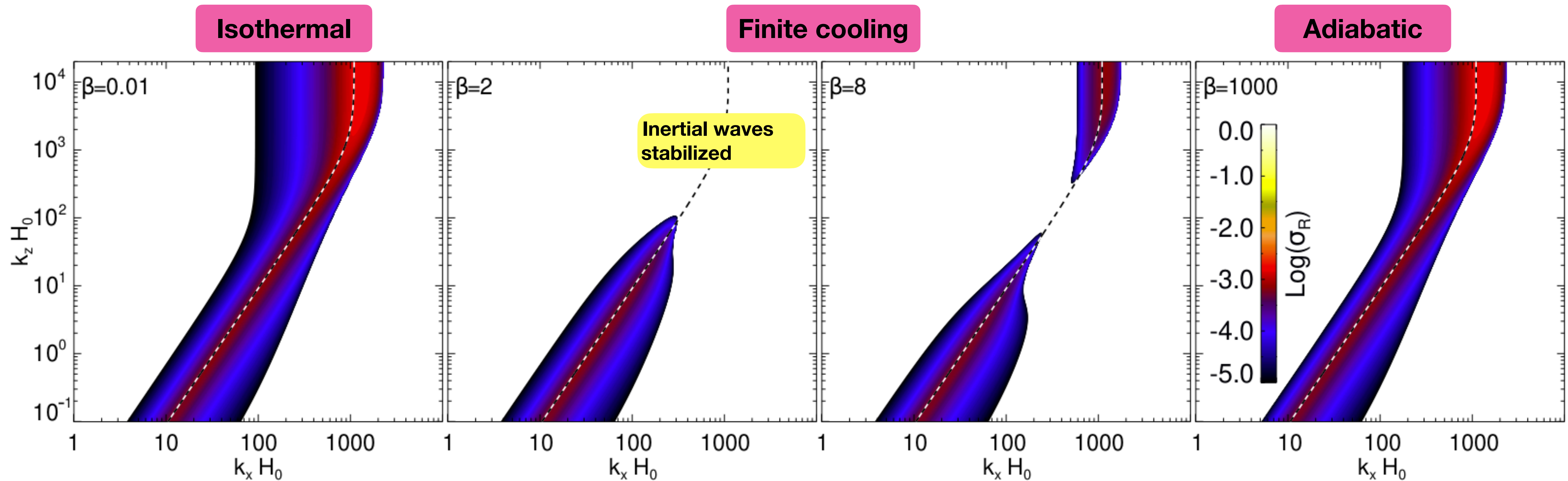
No heating or cooling



(Ishitsu et al, 2009)

**dust layer
dispersed**

SI + radial buoyancy $N_R^2 > 0$



So far, not so good

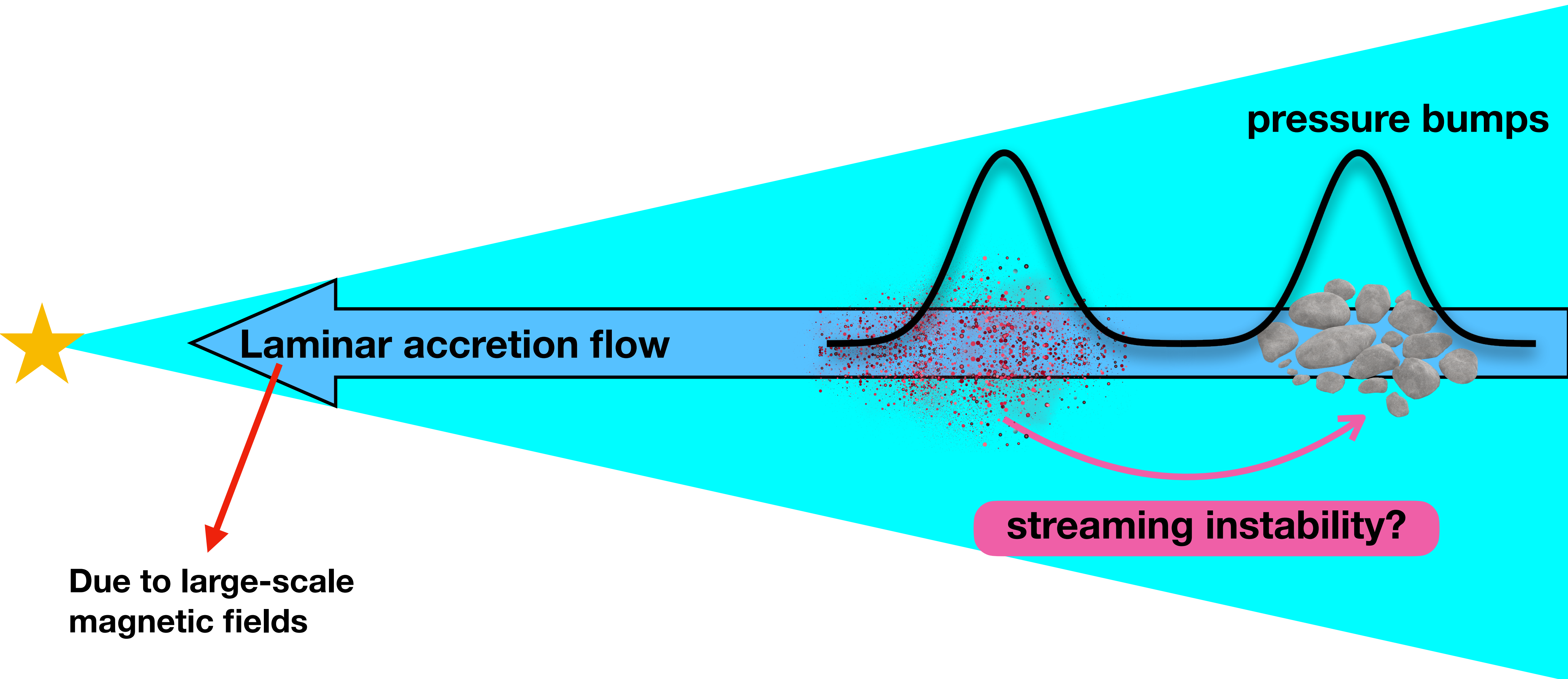
- **turbulence**
- **vertical structure**
- **thermodynamics**



SI weakened

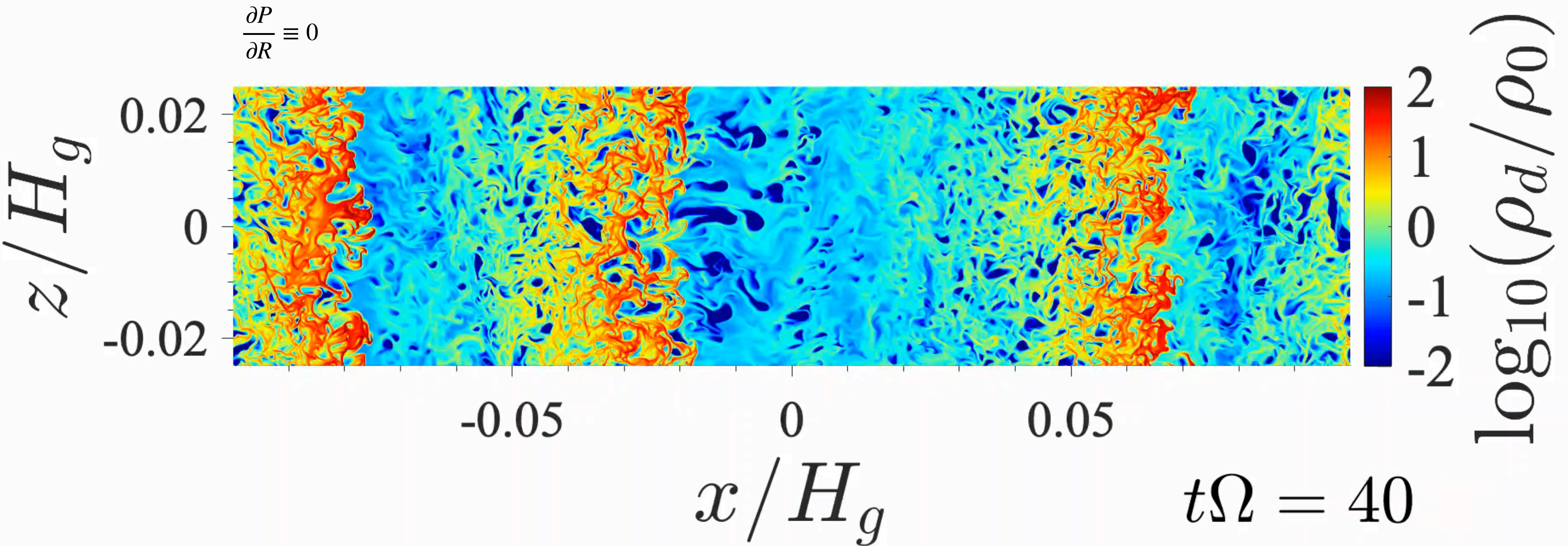


What about magnetic fields?

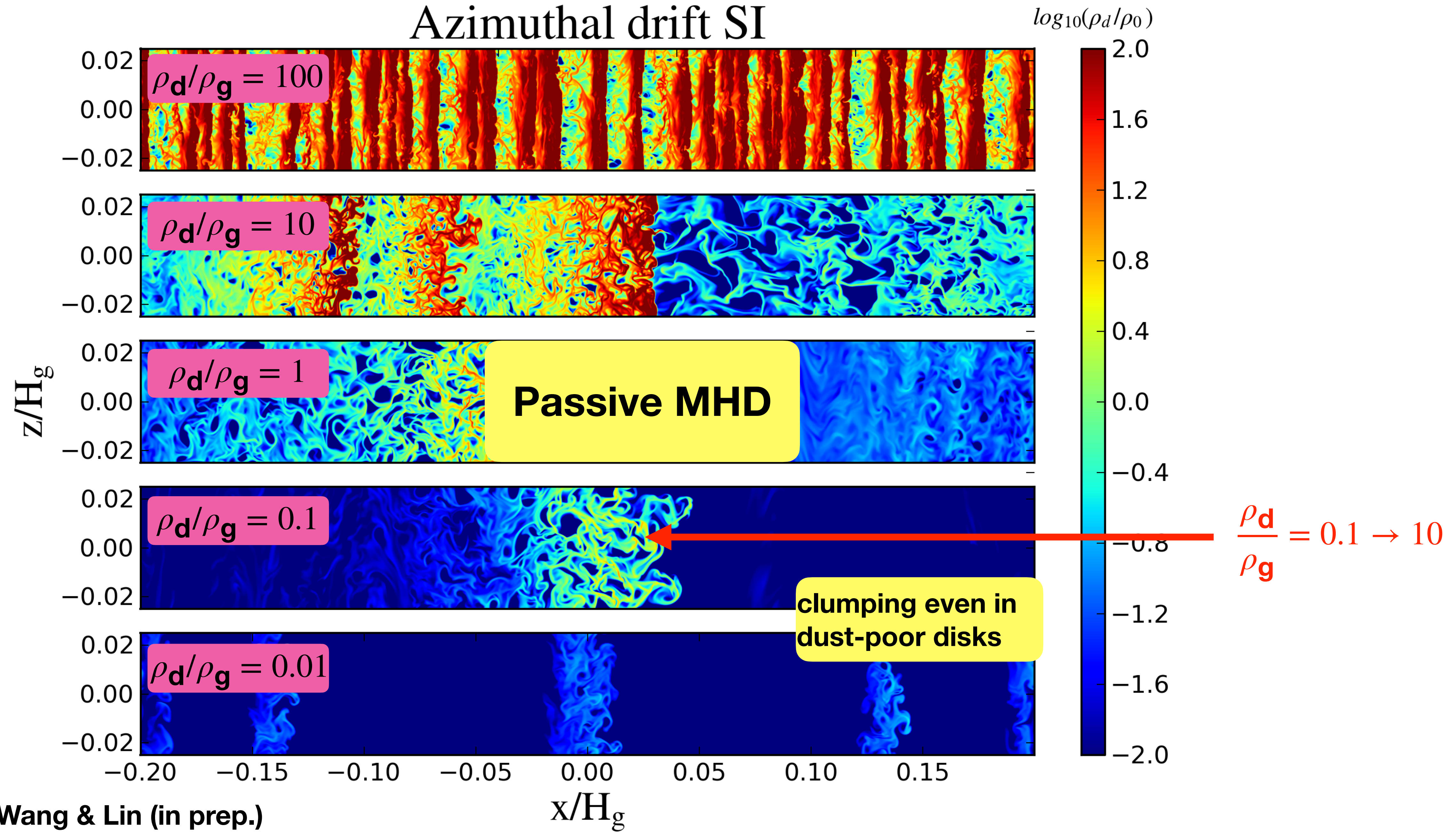


(e.g. Riols et al. 2020, Cui & Bai 2021)

Nonlinear evolution of the SI in accreting disks

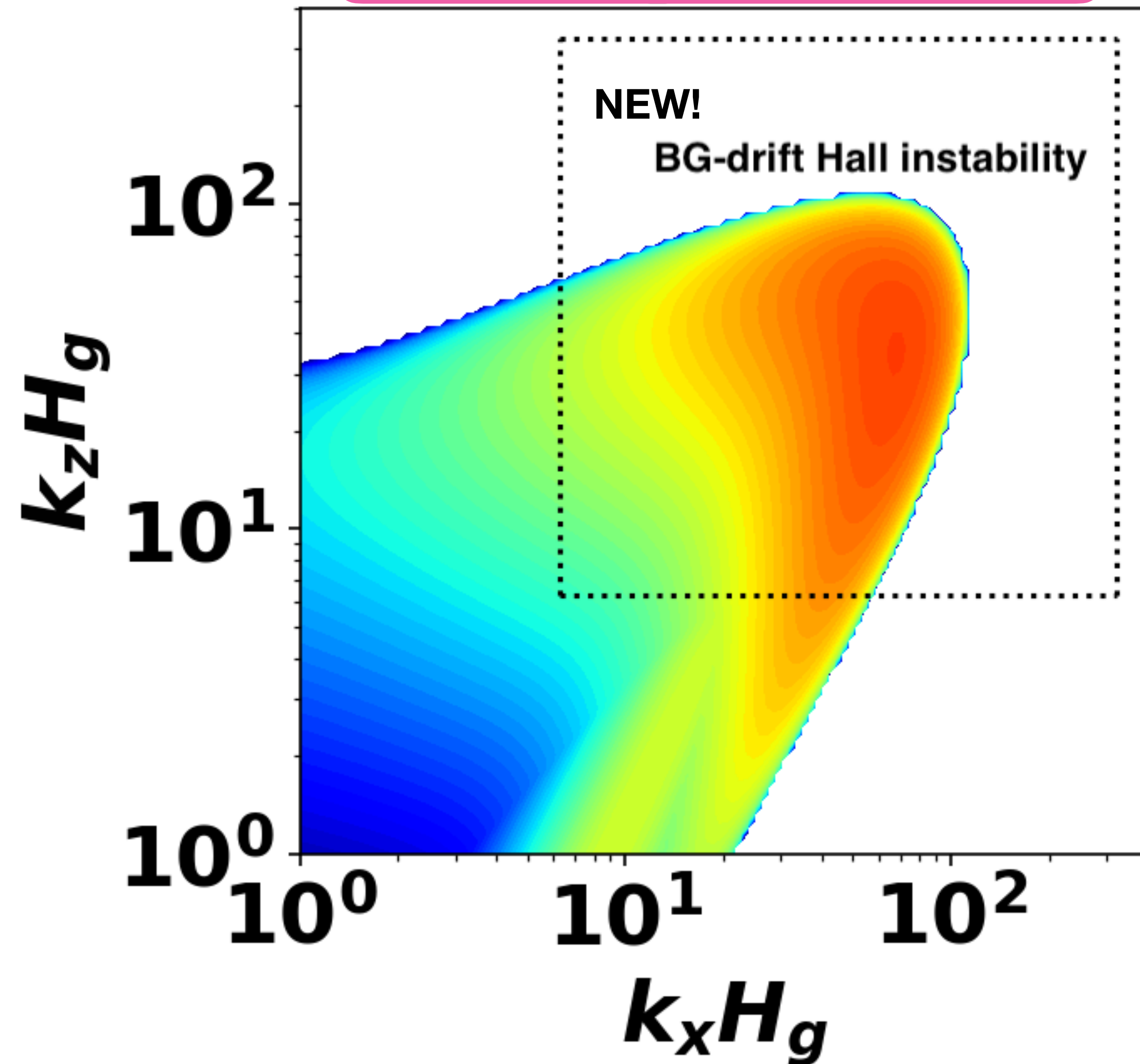


Parameter study

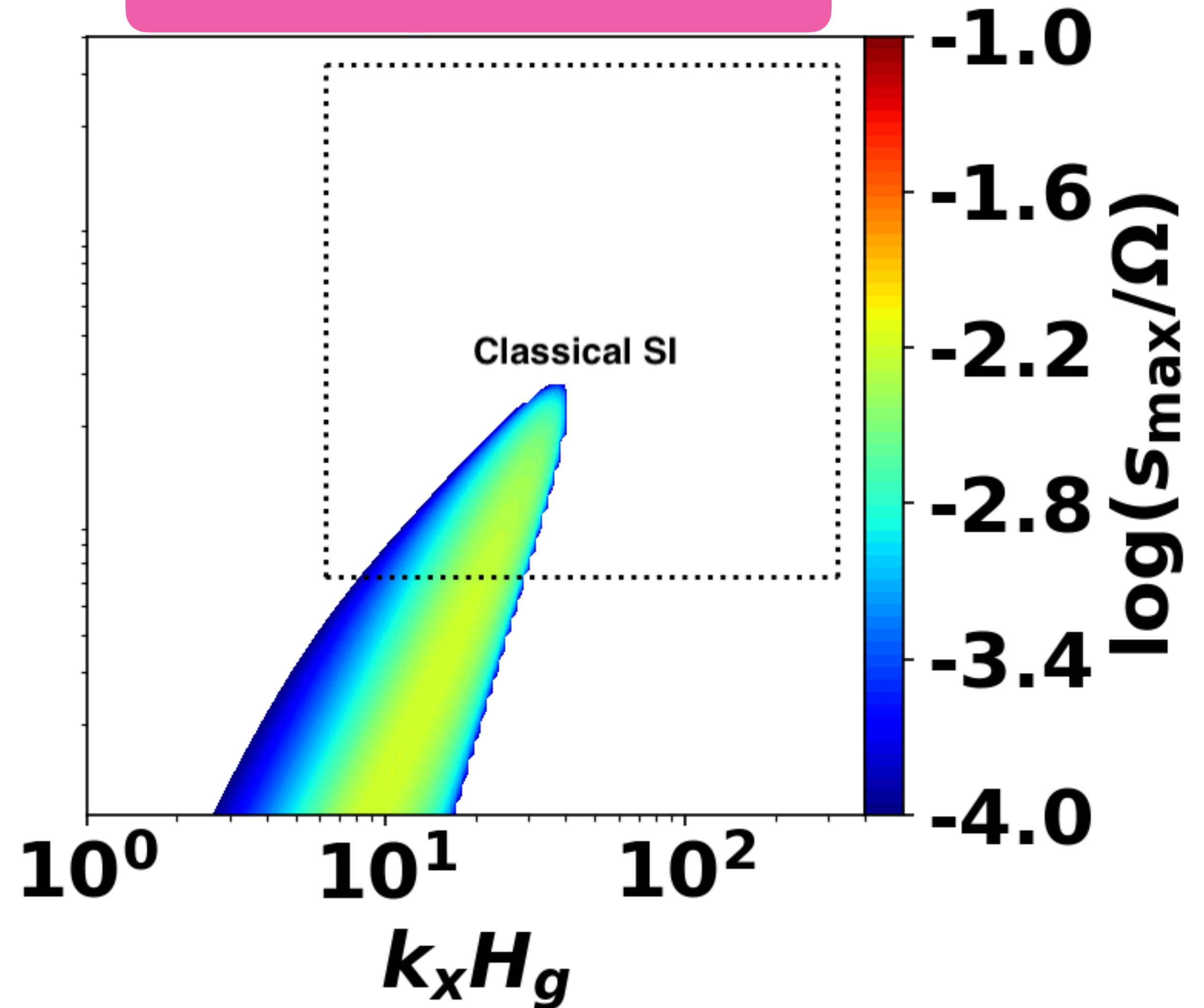


Live MHD & dust dynamics

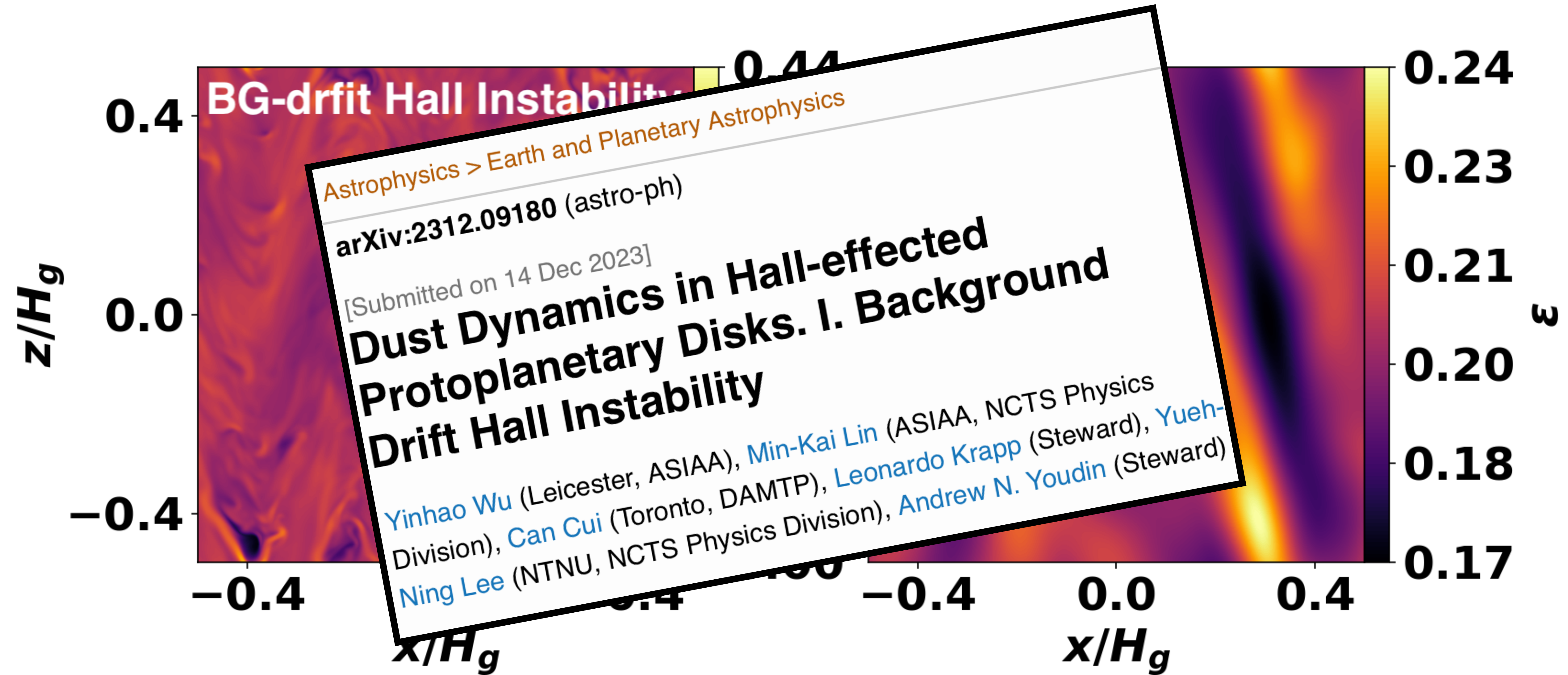
+Hall effect



No MHD



Spectral simulations with Dedalus



Summary

- **Kill the SI:** turbulence, vertical structure, thermodynamics
- **Revive the SI:** gas accretion, magnetic fields
- **Future:** convection, coagulation, improved simulations, etc.

Thank you
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